Environmental Chemistry

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Overview

Lecture	Торіс
1	Introduction / Organisation
2	Thermodynamic
3	Thermodynamic
4	Chemical Equilibrium
5	Photochemistry
6	Photolysis rate, frequency
7	Chemical lifetime
8	Heterogenous Reactions
9	Chemistry of the upper atmosphere
10	Chemistry of the stratosphere
11	Ozone Hole
12	Tropospheric chemistry
13	Tropospheric chemistry
14	Tropospheric chemistry

Organisation:

Exercises: start next week

1st Ex.-Paper this week via e-mail

deadline: Monday or Tuesday next week

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Literature:

P.W.Atkins, Physical Chemistry, Oxford University Press, 1990

Colin Baird, Environmental Chemistry, Free-man and Company, New York, 1995

Guy Brasseur and Susan Solom, Aeronomy of the Middle Atmosphere, D. Reidel Publishing Company, 1986

Guy P. Brassuer, John J. Orlando, Geoffrey S. Tyndall (Eds): Atmospheric Chemistry and Global Change, Oxford University Press, 1999

Finlayson-Pitts B. J. and J. N. Pitts, Atmospheric Chemistry

Richard P. Wayne, Chemistry of Atmospheres, Oxford university Press, 1991

Chemical constituents of the Atmosphere

- Protons and electrons: come from the Sun and created in the atmosphere
- Ion: any species with/without electron
- Atoms: the smallest elemental molecules
- Free radicals: very reactive particles and have at least 1 unpaired electron
- Free radical ions: has electric charge, produce electrical field and can pull other molecules from the distance
- Particles
- Clusters
- Aerosols
- Cloud droplets

The evolution of the atmosphere

The atmosphere compositions were created from the earliest processes in the formation of the solar system and the planets.

The Big Bang created the Universe ~15*10⁹ years

The solar system is 4.6*10⁹ years (and the Earth was formed), based on the radioactive dating of meteorites.

The evolution of Earth see from the geological table:

KANSAS GEOLOGIC TIMETABLE

(Not scaled for geologic time or thickness of deposits)

ERAS	PERIODS	EPOCHS	EST. LENGTH (YEARS)"	DESCRIPTION				
CENOZOIC	QUATERNARY	PLEISTOCENE	10,000+ 1,590,000	Early, the land was stable with some erosion. Glaciers moved into the northeast at least twice. Later the climate was dry. Sand dunes were formed by wind in the west. Volcanic ash was blown in from California, New Mexico, and Wyoming.		Plio-Pleistozan		
	TERTIARY	PLIOCENE MIOCENE OLIGOCENE EOCENE PALEOCENE	3,700,000 18,400,000 12,900,000 21,200,000 6,600,000	Rocks found are part of the Ogallala Formation (sand, gravel, and porous rock), which contains a large quantity of ground water and occurs only in the western third of the state. No rocks were formed in eastern Kansas.	- 1.6	Miozän		
zoic	CRETACEOUS 77,600,000		77,600,000	Much of the western half was covered by seas. Limestone, sandstone, and chalk formed from sea deposits. Fossils can be found in these rocks, which crop out in central and western Kansas.	- 00.4		24	
MESC	JURASSIC		64,000,000	Most rock in Kansas is underground in the west. A few small outcrops are found in the southwest corner. No rocks have been found in Kansas.	Oligozän	Oligozän		
PALEOZOIC	PERMIAN		41,000,000	Much of Kansas was covered by several seas. As they rose and fell, limestone, shale, and chert were deposited. The Flint Hills were formed. When the seas dried up, salt and gypsum were left behind. Salt, now underground, is mined in central Kansas. The Red Hills were formed from deposits of shale, sitistone, sandstone, gypsum, and dolomite.	- 245 MILLION YE	Eozän	- 38 ₪	
	oc	PENNSYLVA		34,000,000	For much of the period the land was flat. Seas and swamps came and went, coal formed in swamps from dead plants. Shale, limestone, sandstone, chert, and conglomerates were deposited. Two ridges of hills, the Nernaha uplift and the Central Kansas uplift, appeared, both are now buried. Pennsylvanian rocks are found at the surface in eastern Kansas.	ARS PAST	Paläozän	— 55 I
	MISSISSIPPI	AN	40,000,000	Repeated layers of limestone, shale, and sandstone indicate that seas rose and fell. Mississippian rocks are the oldest found at the surface and are in the southeast corner; elsewhere these rocks are only underground.	- 320 - 360 - 406 Oberkreide		65	
	DEVONIAN		48,000,000	Seas covered Kansas during much of the period. Limestone, shale, and sandstone deposits are only underground.				
	ORDOVICIAN		30,000,000 67,000,000	deposits are found only underground Seas covered parts of Kansas during much of the period. Dolomite and sandstone are only underground.	- 438			
	CAMBRIAN	(%	65,000,000	Early, the climate was dry and many rocks eroded. Later, parts of Kansas were covered by seas. Dolornite, sandstone, imestone, and shale are now underground.	- 505			
PREG	CAMBRIAN		3,930,000,000	These rocks are the oldest on earth. In Kansas, they are only found deep below the surface and not much is known about them. Many are igneous and metamorphic and have gone through many changes.	- 543		97 I	

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Evolution and changes in atmosphere and climates were modified by the



The changing of population and standard of life lead to the $\rm CO_2$ rising as shown below



The evolution of the atmosphere

Sun system is ~ $4,6 \times 10^9$ years old

At the beginning only a solar haze exists (a cloud with gas, particles and ice)

This cloud was compressed including temperature increase (from 2-3 K to 2000 – 3000 K at the end of compression)

 \rightarrow all compounds evaporated

Inner planets: Venus, Earth, Mars only H_2 , He, C, N_2 and oxygen combined with Si, Fe, Al and Ca

after 10⁷ years planets exist but without atmosphere

The evolution of the atmosphere

Primordialatmosphere

Oxygen was combined \rightarrow so first a reduction of gases in the 1st atmosphere

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1<sup>st</sup> atmosphere: CH<sub>4</sub>, NH<sub>3</sub>, H<sub>2</sub>, H<sub>2</sub>O
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(in the liquid planets: NiO, FeO to Ni and Fe reduced)

 \rightarrow earth's core

Consequently: O₂ was produced

 \rightarrow water vapor, CO_2, N_2 (probably similar to the atmosphere of the vulcanoes today)

H ₂ O	80%
CO ₂	10%
H ₂ S	7%
CO	0.5%
CH ₄	In traces
NH ₃	In traces

Atmosphere today:

Production of:

Hydrosphere: H₂O

Lithosphere: carbonates

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Biosphere: (O<sub>2</sub> was produced)
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2 \text{ CO}_2 + \text{UV-radiation} \rightarrow \text{CO} + \text{O}_2
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2 H<sub>2</sub>O + UV-radiation \rightarrow 2 H<sub>2</sub> + O<sub>2</sub>
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6 \text{ CO}_2 + 6 \text{ H}_2\text{O} + hv + \text{Chlorophyll}\_\rightarrow\_\text{C}_6\text{H}_{12}\text{O}_6 + 6 \text{ O}_2
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Atmosphere – a thin cover



~ 98% of total atmospheric air masses within the lower 30 km

Atmosphere today:



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Ideal Gas

Assumptions:

- · ensemble of individual molecules
- \cdot no interaction apart from collision
- · no chemical reactions
- \cdot no appreciable volume of individual molecules

State properties of a gas: *p*, *T*, *V*, and *n*

Equation of state for the ideal gas: pV = nRT

or
$$pV = NkT$$

p = pressure, V = volume, n = number of moles, N = number of molecules, R = universal gas constant, k = Boltzman constant, T = temperature

Pressure:

1 Pa = 1 N/m² 1 bar = 100 kPa 1 atm = 101.325 kPa = 760 Torr

Pressure of atmosphere is measured with a barometer

Temperature:

- Temperature tells us about the flow of energy
- energy flows from A to B: A has higher temperature
- Energy doesn't flow: A and B have the same temperature (thermal equilibrium)

Thermodynamic temperature scale: $0^{\circ}C = 273.15K$

Gas constant R:

R = 8.314 J K⁻¹ mol⁻¹

Boyle's law: If n, T = constant (isotherm)

> pV = constant $p_2V_2 = p_1V_1$

Gay-Lussac's law:

If n, p or if n, V = constant (isobar, isochor)

 $V_1/V_2 = T_1/T_2$ $p_1/p_2 = T_1/T_2$

Ideal Gas

All gases act as ideal gases at very low pressure;

to good approximation, gases in the atmosphere can be treated as ideal gases with the exception of water vapour (phase changes)

Mixtures of Ideal Gases Dalton's law:

The pressure exerted by a mixture of perfect gases is the sum of the pressures exerted by the individual gases occupying the same volume alone or

$$p = p_A + p_B$$

(air: N₂: 78.08%, O₂: 20.95%, CO₂: 0.04%. These add up to 100% of dry air, so there are almost no other gases (really, 0.002%)(.

Mole fraction:

The mole fraction of a gas X in a mixture is the number of moles of X molecules present (n_X) as a fraction of the total number of moles of molecules (n) in the sample:

$$x_x = \frac{n_x}{n}$$
$$n = n_A + n_B + n_C + \dots$$

with

Partial pressure:

The partial pressure of a gas in a mixture is *defined* as the product of mole fraction of this gas and total pressure of the gas:

$$p_x = x_x p$$



Units

Quantity	Name	Units
Number of molecules	Ν	mol = 6.022 x10 ²³
Density	Ν	particles / m ³
Mass	ρ	kg / m³
Volume mixing ratio	μ	ppmV = 10 ⁻⁶ ppbV = 10 ⁻⁹ pptV = 10 ⁻¹²
Mass Mixing ratio	μ	ppmm =10 ⁻⁶ ppbm =10 ⁻⁹ pptm = 10 ⁻¹²
Different column definition		molec/cm ² DU = 10 ⁻³ cm at STP

Excursion: Ideal Gas Law

The ideal gas law can be expressed in molecules or moles:

$$pV = nRT$$
 or $pV = NkT$

R or R* = Universal gas constant = $8.314 \text{ J} \text{ mol}^{-1} \text{ K}^{-1}$ k = Boltzmann's constant = $1.381 \text{ } 10^{-23} \text{ J} \text{ mol}\text{ec}^{-1} \text{ K}^{-1}$ N_A = Avogadro's number = $6.022 \text{ } 10^{23} \text{ mol}^{-1}$

т

n

Molar quantity:

$$n = \frac{N}{N_A}$$
$$V = \frac{V}{n} \qquad \qquad M =$$

$$R = N_A k$$

Number density:

$$\frac{N}{V} = \frac{p}{kT}$$

Mass density:

$$\rho = \frac{m}{V} = \frac{Nm_x}{V} = \frac{m_x p}{kT} = \frac{M_x p}{N_A kT} = \frac{M_x p}{RT}$$

 m_x = Mass of one molecule x, M_x molar mass of x

In some books, individual gas constants are defined for each gas:

$$MR_x = \frac{R}{M_x}$$

Real Gas:

Interactions between molecules are taken into account

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Moving in a Volume (V) - smaller volume (V - nb)
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With nb: total volume taken by the molecules themselves

Pressure depends on frequency of collisions with walls and the force of each collision - reduction of pressure -a $(n/V)^2$

$$p = \frac{nRT}{V - nb} - a\left(\frac{n}{V}\right)^2$$

Often written like $V_m = V/n$ where $a/V_m^2 =$ internal pressure

$$p = \frac{RT}{V_m - b} - \frac{a}{V_m^2}$$
$$\left(p + \frac{an^2}{V^2}\right)(V - nb) = nRT$$

Ideal gases	Real Gases		
Ideal gases obey all gas laws	Real gases obey gas laws only		
under all conditions of	at low pressures and high		
temperature and pressure.	temperature.		
The volume occupied by the	The volume occupied by the		
molecules is negligible as	molecules is not negligible as		
compared to the total volume	compared to the total volume of		
occupied by the gas.	the gas.		
The force of attraction among	The force of attraction are not		
the molecules are negligible.	negligible at all temperatures		
	and pressures.		
Obeys ideal gas equation	Obeys Van der Waals equation		
PV=nRT	$\left(P + \frac{an^2}{V^2}\right)$ (V-nb) = nRT		

Work, heat, energy are basics of thermodynamic, the most fundamental is work

Work is done by the system to bring a change

Energy is the capacity of a system to do work

Energy changes as a result of temperature difference - energy has transferred as heat

Adiabatic: walls do not permit energy transfer

Exothermic: processes realising energy as heat

Endothermic: processes absorbing energy as heat



1st Law of Thermodynamic

example: increase the temperature of 1kg H_2O from 14°C to 15°C, the work of 4.18 kJ has to be carried out

no energy is lost, it is only exchanged from one to the other (work – heat)

in the case heat is coming to a system \rightarrow variation in volume, pressure, content, temperature = internal energy (total energy of a system) is changing Each system has an internal energy U, which is only changing when heat (Q) or work is changing

$$U_{2} - U_{1} = W + Q = -pdV + Q = \Delta U = \int dU$$

W can be electrically or mechanically

for V = constant (isochor) dV = 0

$$(dU)_V = dQ$$

as follows:

$$(U_2 - U_1)_V = Q$$

Adiabatic process:

$$W_{ad} = U_2 - U_1 = \Delta U$$

Thermodynamic



Mechanical work:

work (compression or expansion) is done by moving an object a distance dz against a force - internal energy will decrease (negative value)

$$dW = -Fdz$$

$$W = -\int_{z_1}^{z_2} F dz = -(z_2 - z_1) = -mgh$$

g: 9.81 m s⁻²

Compression, expansion:

isothermal

$$W = -\int_{V_1}^{V_2} p dV = -nRT \int_{V_1}^{V_2} \frac{dV}{V} = -nRT \ln \frac{V_2}{V_1}$$

Internal energy, heat, work heat have the same unit: Joule (J)

 $1J = 1 \text{ kg m}^2 \text{ s}^{-2}$

1st law: U of a system is constant unless it is changed by doing work or by heating

in a closed system the variation of the internal energy is equal to the exchange of heating with the environment when dV = 0

if dp = 0 (isobar)

$$U_{2} - U_{1} = Q - p(V_{2} - V_{1})$$
$$(U_{2} + pV_{2}) - (U_{1} + pV_{1}) = Q$$

as follows: the exchange of heat is not now the variation of internal energy it is

$$U + pV = H$$

H = Enthalpy

$$(H_2 - H_1) = Q$$

$$H = f(T, p, n_1, n_2, \dots n_k)$$

differential

$$dH = \left(\frac{\partial H}{\partial T}\right)_{p,n} dT + \left(\frac{\partial H}{\partial p}\right)_{T,n} dp + \left(\frac{\partial H}{\partial n_1}\right)_{T,p,n_j} dn_1 + \dots$$
$$dU = \delta Q + \delta W = \delta Q - p dV = \left(\frac{\partial U}{\partial T}\right)_V dT + \left(\frac{\partial U}{\partial V}\right)_T dV$$

$$dH = d\left(U + pV\right) = dU + pdV + Vdp = \left(\frac{\partial H}{\partial T}\right)_V dT + \left(\frac{\partial H}{\partial p}\right)_T dp$$

molar heat capacity for isochor and isobar work:

$$c_V = \left(\frac{\partial Q}{\partial T}\right)_V = \left(\frac{\partial U}{\partial T}\right)_V$$

$$c_{p} = \left(\frac{\partial Q}{\partial T}\right)_{p} = \left(\frac{\partial H}{\partial T}\right)_{p}$$

relation between c_v and c_p :

$$c_p - c_V = \left(\frac{\partial H}{\partial T}\right)_p - \left(\frac{\partial U}{\partial T}\right)_p$$

$$H = U + pV = U + nRT$$

$$c_p - c_V = \left(\frac{\partial U}{\partial T}\right)_p + nR - \left(\frac{\partial U}{\partial T}\right)_p = nR$$

work of adiabatic expansion: as

diabatic expansion: as
$$c_V = \frac{dU}{dT}$$

 $W = \int dU = \int c_V dT = c_V (T_2 - T_1) = c_V \Delta T$

Special cases:

irreversible adiabatic expansion:

$$\Delta T = \frac{-p_{ex}\Delta V}{c_V} \qquad pV = nRT$$

irreversible adiabatic expansion:

$$c_V dT = -pdV \Longrightarrow \frac{c_V dT}{T} = \frac{-nRdV}{V}$$
it is impossible that heat is converted to mechanical energy (work) and we have to find solutions for irreversible processes as follows:



$$\frac{dQ_{rev}}{T} = dS = S_2 - S_1 = nR \ln(\frac{V_2}{V_1})$$

$$dS = \frac{dQ_{rev}}{T} = \left(\frac{\partial S}{\partial T}\right)_V dT + \left(\frac{\partial S}{\partial V}\right)_T dV \quad \text{dependent of T and V}$$

$$dS = \frac{dQ_{rev}}{T} = \left(\frac{\partial S}{\partial T}\right)_p dT + \left(\frac{\partial S}{\partial p}\right)_T dp$$

dependent of T and p

Reversible exchange of heat with environment

$$dS = \frac{dU + pdV}{T} = \frac{1}{T} \left[\left(\frac{\partial U}{\partial T} \right)_V dT + \left(\frac{\partial U}{\partial V} \right)_T dV + pdV \right]$$
$$dS = \frac{dH - Vdp}{T} = \frac{1}{T} \left[\left(\frac{\partial H}{\partial T} \right)_p dT + \left(\frac{\partial H}{\partial p} \right)_T dp - Vdp \right]$$

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$$dS = \frac{dQ_{rev}}{T} = \left(\frac{\partial S}{\partial T}\right)_p dT + \left(\frac{\partial S}{\partial p}\right)_T dp = \frac{dU + pdV}{T} = \frac{dH - Vdp}{T}$$

As follows:

$$dS = \frac{c_V}{T}dT + \frac{1}{T} \left[\left(\frac{\partial U}{\partial V} \right)_T + p \right] dV$$
$$dS = \frac{c_p}{T}dT + \frac{1}{T} \left[\left(\frac{\partial H}{\partial p} \right)_T - V \right] dp$$

in the case of changing the phase (e.g. liquid – solid) reversible: entrophy of changing is used

$$dS = \frac{dQ}{T} = \frac{dH - Vdp}{T}$$

For isobar and isotherm:

$$\Delta S = \frac{\Delta H}{T}$$

Normally all reactions are irreversible but in the case of equilibrium - changing the phase:

$$\Delta S = \frac{\Delta H}{T} \Longrightarrow TdS = dH$$
$$dS = 0$$

in the case of a closed system and using the 1st law of thermodynamic:

$$dS - \frac{dU}{T} \ge 0 \Rightarrow \quad (V, T = \text{constant})$$

 $dS - \frac{dH}{T} \ge 0 \Rightarrow \quad (p, T = \text{constant})$

now with Helmholtz (free energy A) and Gibbs (free enthaphy G) new functions are introduced:

$$A = U - TS$$
$$G = H - TS$$
$$\Delta A = \Delta U - T\Delta S$$

$$\Delta G = \Delta H - T \Delta S$$

in the case G is positiv the reaction is not spontaneous and in contrast if G is negativ the reaction is spontaneous

4 thermodynamic functions: U, H, A, G Closed system, homogenic compound Work only as so called "volume-work"

Internal energy following the 1st law and replacing dQ with entropy and dW with "volume-work":

$$dU = dQ + dW = TdS - pdV$$

Enthalphy: H = U + pV, dH = dU + pdV + Vdp

$$dH = TdS + Vdp$$

dA = dU - TdS - SdT

$$dA = -SdT - pdV$$
$$dG = -SdT + Vdp$$

⇒Guggenheim scheme:



$$T = \left(\frac{\partial U}{\partial S}\right)_{V} = \left(\frac{\partial H}{\partial S}\right)_{p}$$

$$p = -\left(\frac{\partial U}{\partial V}\right)_{S} = -\left(\frac{\partial A}{\partial V}\right)_{T}$$

$$S = -\left(\frac{\partial A}{\partial T}\right)_{V} = -\left(\frac{\partial G}{\partial T}\right)_{P}$$

$$V = \left(\frac{\partial H}{\partial p}\right)_{S} = \left(\frac{\partial G}{\partial p}\right)_{T}$$

Definitions:

Adiabatic: no exchange of energy (walls do not permit energy transfer)
Diathermic: contrast to adiabatic
Exothermic: process when heat (energy) is released
Endothermic: contrast to exothermic (absorbing of energy)

Thermodynamics of mixture

Partial molar volumes of the components of a mixture

$$V_{j} = \left(\frac{\partial V}{\partial n_{j}}\right)_{p,T,n}$$

n_i are the numbers of moles

the composition of the mixture (e.g. H_2O and C_2H_5OH) is changed by the addition of e.g. dn_A of A)

$$dV = \left(\frac{\partial V}{\partial n_A}\right)_{p,T,n_B} dn_A + \left(\frac{\partial V}{\partial n_B}\right)_{p,T,n_A} dn_B = V_A dn_A + V_B dn_B$$

in the case we know the partial molar volumes then the total volume

$$V = n_A V_A + n_B V_B$$

partial molar quantity can also be expressed by the partial molar Gibbs function

$$\mu_{j} = \left(\frac{\partial G}{\partial n_{j}}\right)_{p,T,n'}$$

 \rightarrow total Gibbs function

$$G = n_A \mu_A + n_B \mu_B$$

 $\mu_{A}\,\mu_{B}$ are the chemical potentials (or molar free enthaphy) at the composition changing

$$dG = n_A d\mu_A + n_B d\mu_B$$

at constant pressure and temperature

$$n_A d\mu_A + n_B d\mu_B = 0$$
$$\sum n_j d\mu_j = 0$$

two gases at a temperature and a pressure

 \rightarrow Gibbs function of mixing

$$\Delta G_{mix} = n_A RT \ln \frac{p_A}{p} + n_B RT \ln \frac{p_B}{p}$$

other thermodynamic mixing functions

$$-S = \frac{\partial G}{\partial T}_{p,n}$$

a mixture of a perfect gas

$$\Delta S_{mix} = -\left(\frac{\partial \Delta G_{mix}}{\partial T}\right)_{p,n_A,n_B} = -nR\left(x_A \ln x_A + x_B \ln x_B\right)$$

since $Inx < 0 \rightarrow \Delta S_{mix} > 0$

the enthalpy of mixing is then:

$$\Delta G = \Delta H - T \Delta S$$

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Chemical equilibrium:

$$\left(\Delta G\right)_{p,T}=0$$

Chemical reaction, if the chemical potential is decreasing:

$$\left(\Delta G\right)_{p,T} < 0$$

For isothermal processes $\Delta T = 0$

and $\Delta H_{mix} = 0$ (constant p, T)

The chemical potential of liquids :

At equilibrium:

$$\mu = \mu^{\Theta} + RT \ln = \left(\frac{p}{p^{\Theta}}\right)$$

with the chemical potential

$$\mu = \left(\frac{\partial G}{\partial n}\right)_{p,T}$$

Ideal solution:

- The chemical potential of pure liquid A is $\mu_{A}^{*}(I)$
- The vapor pessure is then: p_A^*
- Chemical potential in vapor is:

$$\mu_{A^{\Theta}} + RT$$
In

the both chemical potential are equal at equilibrium:

$$\frac{p_{A}}{p^{\Theta}}$$

in addition an other substance is present then:

The chemical potential of A in the liquid is $\mu_A(I)$ and its vapor pessure is then: p_A

In this case:

$$\mu_A(l) = \mu_A^{\Theta} + RT \ln\left(\frac{p_A}{p^{\Theta}}\right)$$

to eliminate the standard potential of the gas the eqn. are combined

Henry's Law

With x_B (mole fraction of solute) and K_B is a constant (with dimensions of pressure) Spontaneous chemical reactions

A + B \rightarrow C + D A, Reaction completed: all the transformed to products B, reaction partially: equal A + B and C + D

C, no reaction: only reactants

 $A \to B$

perfect gas equilibrium:

 $A \leftrightarrow B$ equilibrium

$$\Delta G = \mu_B - \mu_A = \left(\mu_B^{\Theta} + RT\ln\frac{p_B}{p^{\Theta}}\right) - \left(\mu_A^{\Theta} + RT\ln\frac{p_A}{p^{\Theta}}\right)$$
$$\Delta G = \Delta G^{\Theta} + RT\ln\frac{p_B}{p_A}$$

the ratio of the partial pressures is an example of reaction quotient Q_p

$$\Delta G = \Delta G^{\Theta} + RT \ln Q_p$$

at equilibrium $\Delta G = 0$, the value of the reaction quotient = equilibrium constant K_p

$$K_{p} = \left(Q_{p}\right)_{equilibriu\ m} = \left(\frac{p_{B}}{p_{A}}\right)_{eq}$$

then with $\Delta G = 0$, $Q_p = K_p$

$$RT \ln K_p = -\Delta G^{\Theta}$$

e.g
$$2A + 3B \rightarrow C + 2D$$

$$\Delta G^{\Theta} = -2\mu_A^{\Theta} - 3\mu_B^{\Theta} + \mu_C^{\Theta} + 3\mu_D^{\Theta}$$

$$K = (Q)_{equilibrium} = \left(\frac{a_C a_D^2}{a_A^2 a_B^3}\right)_{eq}$$

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Gibbs-Energy is defined:

$$G = U + pV - TS$$
$$H = U + pV \longrightarrow U = H - pV$$

$$G = H - pV + pV - TS = H - TS$$

With U, V, S and H

As

And

Often the differential form is used:

$$dG = dU + d(pV) - d(TS) = dQ + dW + d(pV) - d(TS)$$
$$dS = \frac{dQ}{T} \rightarrow TdS = dQ$$

dG = TdS - pdV + pdV + Vdp - TdS - SdT = Vdp - SdT

If the number of moles are not const:

$$dU = TdS - pdV + \Sigma \mu i dn i$$

$$dG = -SdT + Vdp + \Sigma \mu i dn i$$

At the equilibrium, the term of Gibbs energy equal zero, .

$$dG = 0$$

Any reactions try to move toward the equilibrium in both reactant and product and time is needed to reach the equilibrium e.g. in the atmosphere at the present time, CO_2 is rising to reach the equilibrium temperature.

Chemical Equilibrium: the state in which the chemical activities or concentrations of the reactants and products have no net changed over time. The reaction rates of the forward and reverse reactions are generally not zero but, being equal, there are no net changes in any of the reactant or product concentrations. This process is called dynamical eugilibrium

In this path, we will use the thermodynamics to describe the equilibrium under the reaction conditions. In term of Gibbs energy for the perfect (ideal) gas,

$$\Delta G = \Delta G^{\Theta} + RT \ln \frac{p_A}{p_B}$$

as p_A is the partial pressure of reactant A and p_B is the partial pressure of product B.

At the equilibrium, the term of Gibbs energy equal zero, .



Let consider the reaction;

$$v_A A + v_B B \Leftrightarrow v_C C + v_D D$$

in term of the reaction activity, we define the ratio of partial pressures at equilibrium as the thermodynamic equilibrium constant (as denoted K_p). Thus the Gibbs energy of the standard state can be written as:

$$-\Delta G^{\Theta} = RT \ln K_p \qquad \qquad K_p = \frac{a_C^{\nu_C} a_D^{\nu_D}}{a_A^{\nu_A} a_B^{\nu_B}} = \prod_J a^{\nu_J} - \Delta G^{\Theta} / \dot{R}T = \ln K_p$$

 $a^{\nu_i}J$ is the partial pressure of substance J having mole. ν_i The equilibrium constant can be defined in term of concentration, K_c.

$$K_{C} = \frac{\left[D\right]^{\nu_{D}} \left[C\right]^{\nu_{C}}}{\left[A\right]^{\nu_{A}} \left[B\right]^{\nu_{B}}}$$

Example: combustion process in motor:

$$N_2 + O_2 \Leftrightarrow 2NO$$

$$K = \frac{k_{forward}}{k_{backward}} = \frac{[NO]^2}{[N_2][O_2]}$$

The rate constants and the concentration of the reactants and products are considered in the equilibrium constant

The relationship between equilibrium constants in terms of partial pressure (K_P) and concentration (K_C) can be determined by using the ideal gas law.

$$p = \frac{nRT}{V}$$

$$K_{P} = \frac{a_{C}^{\nu_{C}} a_{D}^{\nu_{D}}}{a_{A}^{\nu_{A}} a_{B}^{\nu_{B}}} = \frac{\left(\frac{\nu_{C}}{V} RT\right)^{\nu_{C}} \left(\frac{\nu_{D}}{V} RT\right)^{\nu_{D}}}{\left(\frac{\nu_{A}}{V} RT\right)^{\nu_{A}} \left(\frac{\nu_{B}}{V} RT\right)^{\nu_{B}}} = \frac{\left[C\right]^{\nu_{C}} \left[D\right]^{\nu_{D}}}{\left[A\right]^{\nu_{A}} \left[B\right]^{\nu_{B}}} \left(RT\right)^{\Delta\nu_{j}}$$

$$K_P = K_C (RT)^{\Delta v_j}$$

How equilibria respond to pressure?

Henry's law (real solution)

For the real solutions at low concentration, the proportional of its mole fraction (X_B) and the vapour pressure of the pure substance (P_B) is constant (K_B).

$$p_B = X_B K_B$$

Or the concentration of a gas in a solution is direct proportional to the partial pressure of the gas over the solution. The propornality is difined by K (Henry's law coefficients representing solubility (i.e. aqueous-phase composition divided by gas-phase composition). This law is mostly used to understand the real atmosphere, for example the reaction:

$$N_2O_5(g) + H_2O(I) \rightarrow 2HNO_3$$

in form of stick to N_2O_5 (g) stick to the surface of H_2O (I)

Additional superscripts denote which properties are used for Henry's law:

symbol	definition	description	unit(s)
k _H cp	c/p	aq-concentration / partial pressure	[mol/(m ³ *Pa)] or [M/atm]
k _H cc	c/c	aq-concentration / gas-concentration	dimensionless
k _H ^{bp}	b/p	molality / partial pressure	[mol/(kg*Pa)] or [mol/(kg*atm)]
k _H ^{xp}	x/p	molar fraction / partial pressure	[1/atm]
k _{H,inv} px	p/x	partial pressure / molar fraction	[atm]
k _{H,inv} pc	p/c	partial pressure / aq-concentration	[m³*Pa/mol]
k _{H,inv} cc	c/c	gas-concentration / aq-concentration	dimensionless

Raoult's law (ideal solution)

The ratio of the partial pressure of each component to its vapour pressure as a pure liquid

$$\frac{p_A}{p_A^{*}}$$

is approximately equal to the mole fraction of A in the liquid mixture. p_A : partial pressure over liquid, p_{A^*} : partial pressure of liquid

$$p_A = X_A p_A *$$

Example:

$$A \Leftrightarrow 2B$$

The reaction:

at t=0 the molar reactant A = n and the product B = 0

at t=t the molar reactant A = $n(1-\alpha)$ and the product B = $2n\alpha$

the mole fraction
$$x_A = \frac{(1-\alpha)n}{(1-\alpha)n + 2n\alpha} = \frac{1-\alpha}{1+\alpha}$$

the mole fraction
$$x_B = \frac{2n\alpha}{(1-\alpha)n + 2n\alpha} = \frac{2n}{1+\alpha}$$

$$K_{p} = \frac{\left(\frac{p_{B}}{p_{\Theta}}\right)^{2}}{\left(\frac{p_{A}}{p_{\Theta}}\right)^{2}} = \frac{\left(\frac{x_{B}p}{p_{\Theta}}\right)^{2}}{\left(\frac{x_{A}p}{p_{\Theta}}\right)^{2}} = \frac{x_{B}^{2}p}{x_{A}p_{\Theta}} = \frac{4\alpha^{2}p}{(1-\alpha)p_{\Theta}}$$
$$\alpha = \left(\frac{K_{p}}{K_{p}+4\frac{p}{p_{\Theta}}}\right)^{\frac{1}{2}}$$

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It shows that the amounts of A and B depend on pressure. For example, if P is increased, the amount of α will decrease to reach the new equilibrium, (the equilibrium constant, K_p, is independent of pressure.

$$\left(\frac{\partial K}{\partial p}\right)_T = 0$$

This corresponds to Le Chetalier's principle.

Le Chetalier principle

"A system at equilibrium, when subjected to a disturbance, responds in a way that tends to minimize the effect of the disturbance."

e.g. if you want to have more products \rightarrow change:

- 1. T or/and p
- 2. Increase concentration of reactants or decrease the concentratopn of product continously
- 3. An additional catalysator

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How equilibria respond to temperature?

From the Gibbs energy of the standard state at the equilibrium;

$$-\Delta G^{\Theta} = RT \ln K_p$$
$$\ln K_p = -\frac{\Delta G^{\Theta}}{RT}$$
$$\frac{\partial}{\partial T} \ln K_p = -\frac{1}{R} \frac{\partial}{\partial T} \frac{\Delta G^{\Theta}}{T}$$

Gibbs-Helmholtz-Equation

G = H - TSconnection G, H, TS: $\left(\frac{\partial G}{\partial T}\right) = -S$ Isobare modication of G with T: $\left(\frac{\partial G}{\partial T}\right) = -\frac{H-G}{T} \checkmark$ Replacement of S: $\left(\frac{\partial G}{\partial T}\right)_{n} - \frac{G}{T} = -\frac{H}{T}$ All G on left side: $\frac{1}{T} \left(\frac{\partial G}{\partial T} \right)_{T} - \frac{G}{T^2} = -\frac{H}{T^2}$ Division by *T*:

Gibbs-Helmholtz-Equation

$$\frac{1}{T} \left(\frac{\partial G}{\partial T} \right)_p - \frac{G}{T^2} = -\frac{H}{T^2}$$

 $\frac{\partial}{\partial T} \left(\frac{G}{T} \right)_p = -\frac{H}{T^2} \quad \longleftarrow$

Summary on right side:

Product law:

$$\frac{\partial}{\partial T} \left(\frac{G}{T} \right)_p = \frac{\partial}{\partial T} \left(G \cdot \frac{1}{T} \right)_p = \frac{\partial}{\partial T} \left(G \right)_p \cdot \frac{1}{T} + G \cdot \left(-\frac{1}{T^2} \right) - \frac{1}{T^2} + \frac{1}{T^$$

Gibbs-Helmholtz-Equation

$$\frac{\partial}{\partial T} \left(\frac{G}{T} \right)_p = -\frac{H}{T^2}$$

Connection of temperature dependency of Gibbs-Potential and Enthalpy

$$\frac{\partial}{\partial T} \left(\frac{\Delta G}{T} \right)_p = -\frac{\Delta H}{T^2}$$

van't Hoff' Reaction isobare

$$\left(\frac{\partial \ln K}{\partial T}\right)_p = \frac{\Delta H^\circ}{RT^2}$$

Use the Gibbs-Helmholz equation:

$$\frac{\partial}{\partial} \left(\frac{\Delta G^{\Theta}}{T} \right) = -\frac{\Delta H^{\Theta}}{T^2}$$

Therefore,

$$\frac{\partial \ln K_p}{T^2 \partial (1/T)} = \frac{\Delta H^{\Theta}}{RT^2}$$

$$\frac{\partial \ln K_p}{\partial (1/T)} = \frac{\Delta H^{\Theta}}{R} \qquad \text{and} \qquad \frac{\partial \ln K_p}{\partial T} = \frac{\Delta H^{\Theta}}{RT^2}$$

These are the simplifies of the van 't Hoff equations, which show the changing of temperature in equilibrium is dependence on the changing of enthalpy.
Photochemistry

Photochemisty: chemical reactions initiated by light

Conditions: absorption of light of the molecule which reacts in a wavelenght region of the incoming light

Absorption of light leads to an energetic exited level of a molecule

This additional energy can be lost by emission (singulet-level = flourescence) or triplel-level = phosphorescence)

Or without radiation = collision, chemical reaction









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Lifetime:

crossing from one to the other level, stay on an exited level (eletronic: in VIS-wavelenght region = 10⁻⁹s rotational: in mirco wave region = 10³s)

Photochemistry

Most chemical processes in the atmosphere are initiated by photons. The general reaction:

 $ABC + h\nu \rightarrow product$

The rotational energy At low energy:

in microwave or FIR wavelengths) : $\lambda > 25 \ \mu m$

$$ABC + h\nu \rightarrow ABC *$$
 (rotation excited)

Molecules will be emitted or collided the original molecule itself.

$$ABC + h\nu \rightarrow ABC^* \rightarrow ABC + h\nu$$

or

$$ABC + h\nu \rightarrow ABC^* \rightarrow ABC$$

The rotational energy is E_J :

$$E_J = BJ(J+1)$$

J= constant rotational number

$$B \approx \frac{1}{I}$$
 $I = \mu r^2$

$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$

= reduced mass

For example: $I_{HCI} = 1$ and $I_{CI2} = 1/37$,

The small value of I has large B, it introduces to the large energy E_J value, therefore .

$$E_{J,Cl_2} > E_{J,HCl}$$

The vibrational energy

At higher energy:

in thermal IR wavelength, 2.5 μ m < λ < 25 μ m.

 $ABC + h\nu \rightarrow ABC^*$



(vibrational excited)



 $ABC + h\nu \rightarrow ABC^* \rightarrow ABC + h\nu''$ emit

collision

 $ABC + h\nu \rightarrow ABC^* \rightarrow ABC$

Molecules will be collided or emitted the original molecule itself with the vibrational energy E_v :

$$E_{\nu} = h \nu_0 \left(\nu + \frac{1}{2} \right) \qquad \nu_0 = \frac{1}{2\pi} \sqrt{\frac{k}{\mu}} \qquad \mu = \frac{1}{\mu_1} + \frac{1}{\mu_2}$$

v= constant vibrational number

Combination of rotation and vibration



Combination of rotation and vibration (spectrum)



Spectrum of benzene





Spectrum of CO



Line	$\widetilde{\mathcal{V}}$	$\Delta \widetilde{\nu}$	Line	$\widetilde{\nu}$	$\Delta \widetilde{\nu}$
<i>P</i> (1)	213943		<i>R</i> (0)	214708	
		3.88			3.78
P(2)	213555		<i>R</i> (1)	215086	
		3.92			3.73
P(3)	213163		<i>R</i> (2)	215459	
		3.95			3.72
<i>P</i> (4)	212768		<i>R</i> (3)	215881	
		3.98			3.66
P(5)	212370		<i>R</i> (4)	216197	
			-		

 $\Delta v = v_0 - 2BJ$

 $\Delta v = v_0 + 2B(J+1)$

The electronic energy

At high energy: in UV and visible wavelengths, he molecules will vibrated, rotated and have electronic transition. In the physical processes lead to emission of radiation and energy is converted to heat.

In the photochemical processes lead to the several pathways as shown in the figure.



Fig. 3.1. Pathways for loss of electronic excitation that are of importance in atmospheric chemistry. The use of the symbols * and § illustrates the presence of electronic excitation: the products of any of the processes may be excited. With the exception of pathways (i) and (iv), excited atoms can participate as well as excited molecules.

Atomic and molecular spectroscopy:

In general, an atomic term symbol is of the form ${}^{2S+1}L_{I}$,

where 2S+1 represents the spin multiplicity of the atom,

L represents the orbital angular momentum of the atom and J represents the total quantum number, J = L+S.

Example

Oxygen atom

Electronic configuration is $1s^2 2s^2 2p^4$.

Atomic term symbol at ground state is ${}^{3}P_{2}$.







Oxygen molecule

In atmosphere, the oxygen molecule is formed by photolysis of ozone.





Photolysis rates and frequency

The general reaction:

 J_1

$$AB + h\nu \rightarrow product$$

The photolysis rate from the reaction is

$$\frac{d}{dt}[AB] = -J_1[AB]$$

where as [AB] = concentration of AB substance

$$\frac{d}{dt}[AB]$$
 = photolysis rate

= photolysis frequency (similar to $k \rightarrow \tau = 1/J$ with J in [s⁻¹])

<u>J</u>: Integration of product of spectral actinic flux I (intensity of light), spectral absorption coefficient σ [cm²](area in which one photon is absorbed) and relative quantum yield (\emptyset [cm⁻² s⁻¹])

$$J_1(\lambda,T) = \int_{\lambda_{\min}}^{\lambda_{ax}} I(\lambda)\sigma(\lambda,T)\Phi(\lambda,T)d\lambda$$

Photons can arrive from all directions on the molecule \rightarrow spheric flux has to be considered $I(2) = \int \int R(2, \theta) d\Theta$

$$I(\lambda) = \int_{0}^{1} \int_{0}^{0} R(\lambda,\theta) d\Omega$$

$$d\Omega = \sin\theta d\theta d\phi$$

- dΩ infinitesimal solid angle
- Φ azimuth
- Θ solar zenith angle

Quantum yield (Ø) is the relative efficiency of various photophysical and photochemical processes.

 $\phi = \frac{number of molecule photolysed by the chemical process}{total number of photo absorbed}$

For example:
$$O_2 + h \nu \rightarrow O + O$$

 Φ is wavelength dependent, all important at different λ .



Quantum yields for NO₃ photolysis: (dotted line)

 $NO_3 \rightarrow NO_2 + O$

(solid line) $NO_3 \rightarrow NO + O_2$

(dashed line), fluorescence quantum yield (From F-P&P)

J depends on

- Light intensity from all directions "Actinic flux"
- Absorption cross section (σ)
- Quantum yield for photodissociation (Φ)
- All are functions of wavelength

$$J = \int_{\lambda_1}^{\lambda_2} I_{\lambda} \sigma(\lambda) \phi(\lambda) d\lambda$$

 J_1

e.g. Destruction and production of
$$O_3$$

$$O_3$$
 + hv (I< 310 nm) $\rightarrow O + O_2$

 $O + O_2 + M \rightarrow O_3 + M$ k₃

Productionrate:

 $d [O_3]/dt = k_3 [O][O_2][M]$

Both processes production and destruction are very fast \rightarrow equilibrium:

Production = destruction

 $J_1[O_3] = k_3[O][O_2][M]$

$$[O]/[O_3] = J_1/k_3[O_2][M]$$

As $[O_2]$ and [M] decrease with increasing height \rightarrow $[O]/[O_3]$ increase

In addition O is produced by the photolysis of O₂

$$O_2$$
 + hv (λ < 230 nm) \rightarrow O + O

 $d [O]/dt = J_2 [O_2]$



Measurements at the same time of ozone over Texas



Photolysis rates of O_2 (J_2) and O_3 (J_1) at different heights h.

Kinetics

Order of reaction:

Overall = sum of more than one elementary reaction

The order with respond to the individual reactant

First order reaction

 $A \rightarrow B$

$$-\frac{d[A]}{dt} = k[A]$$

$$-\frac{d[A]}{[A]} = kdt$$



$$\int_{[A]_0}^{[A]} \frac{d[A]}{[A]} = -\int_0^t k dt$$

$$\ln \frac{[A]}{[A]_0} = -kt \longrightarrow [A] = [A]_0 \exp(-kt)$$





1st order: $A \rightarrow C + D$ 2nd order: $A + B \rightarrow C + D$ 3rd order $A + B + M \rightarrow C + M$

Molecularity of Reactions

• Unimolecular: overall is first order.

$$AB \rightarrow A + B$$

• Bimolecular: overall is second order, two body reactions

$$A + BC \to AB + C$$

• Termolecular: overall is third order, three body reactions, association reactions. example

 $O + O_2 + M \rightarrow O_3 + M$

M = usually N_2 or O_2 for atmosphere

The general reaction: $aA + bB \rightarrow cC + dD$

Rate of reaction
$$= \frac{d[A]}{dt} = \frac{d[B]}{dt} = -\frac{d[C]}{dt} = -\frac{d[D]}{dt}$$

Rate of reaction in terms of concentrations can be written:

Rate of reaction, = $k[A]^a[B]^b$ where as k is rate constant.

For elementary reactions (a + b \leq 3), the <u>Arrhenius equation</u> can be apply to calculate the rate constant k

 \rightarrow temperature dependence of rate constant is given by Arrhenius equation



Distance along Reaction Coordinate

Photolysis rate, frequency The collision theory $A + BC \rightarrow [ABC]$ nc Collision rate for molecules A and BC = 4 = number of molecular density (concentration) and where as n С = mean velocity

$$-\frac{d[A]}{dt} = [A][BC]\sigma \bar{c}$$

when no reaction occurs.

Low concentration = Few collisions

High concentration = More collisions

$$\frac{d[A]}{dt} = [A][BC]\sigma \bar{c} \exp\left(-\frac{E_a}{RT}\right) \quad \text{when reaction occurs.}$$
$$\bar{c} = \sqrt{\frac{8kT}{\pi\mu}} \quad \text{k = Boltzmann constant}$$
$$\mu = \text{reduced mass}$$

$$\sigma = \pi (r_A + r_{BC})^2$$
 = collision cross section

Therefore, in collision theory:

$$k_{C} = \sigma \bar{c} \exp\left(-\frac{E_{a}}{RT}\right)$$

Chemical lifetime



Second order reaction:
$$A + BC \rightarrow product$$

Rate of reaction:
$$\frac{d}{dt}[A] = -k_2[A][BC]$$

 $\frac{d}{dt}[A] = k[B]$

The true rate law is forced to be first-order form, called a pseudo-first-order rate law and

$$k = k_2[A]$$

Unit of k:= cm³ molecule⁻¹s⁻¹
Chemical lifetime
$$\tau_1(A) = \frac{1}{k} = \frac{1}{k_2[BC]}$$

 $[OH]_{global} = 5x10^5$ molecule cm⁻³

1st example: $OH + CH_4 \rightarrow CH_3 + H_2O$

 $k_{CH_4}(298K) = 1x10^{-14} \text{ cm}^3 \text{ molecules}^{-1} s^{-1}$ $\tau_{1^{(CH_4)}} = \frac{1}{k_3[OH]} = 2x10^8 s \cong 10 \text{ years}$

2nd example: $OH + CO + M \rightarrow CO_2 + H + M$ $k_{CO}[M] \cong 1x10^{-13} \, cm^3 molecules^{-1}s^{-1}$ $\tau_1(CO) = \frac{1}{k[OH]} \cong 2x10^7 s \cong 8months$

3rd example: $OH + NO_2 + M \rightarrow HNO_3 + M$ $k_{NO_2}[M] \cong 1x10^{-11} cm^3 molecules^{-1}s^{-1}$ $\tau_1(NO_2) = \frac{1}{k[OH]} \cong 2x10^5 s \cong 2.3 days$

In equilibrium (or stationary state or steady state), the production rate (P) and loss rate (L) are equal.

$$\frac{d}{dt}[A] = 0 = P - L$$

Example: For equilibrium reaction:

$$NO_{2} + NO_{3} + M \leftrightarrow N_{2}O_{5} + M$$

$$\frac{d[N_{2}O_{5}]}{dt} = 0 = P - L = k_{+}[NO_{2}][NO_{3}] - k_{-}[N_{2}O_{5}]$$

$$K = \frac{k_{-}}{k_{+}} = \frac{[NO_{2}][NO_{3}]}{[N_{2}O_{5}]}$$
For the termolecular (3rd order) reaction:

$$A + BC \rightarrow ABC^*$$

$$ABC^* \rightarrow A + BC$$
 k₂

$$ABC * + M \rightarrow ABC + M = k_3$$

$$\frac{d[ABC]}{dt} = 0 = k_1[A][BC] - k_2[ABC^*] - k_3[ABC^*][M]$$

Assume that [ABC*] is in the steady state:

$$\frac{d[ABC]}{dt} = k_1[A][BC] - (k_2 + k_3[M])[ABC^*]$$

$$\frac{d[ABC]}{dt} = k_1[A][BC] - (k_2 + k_3[M])[ABC^*]$$
$$[ABC^*] = \frac{k_1[A][BC]}{k_2 + k_3[M]}$$
$$d[ABC] = d[A] - d[B] - k [ABC^*][$$

$$\frac{d[ABC]}{dt} = -\frac{d[A]}{dt} = -\frac{d[B]}{dt} = k_3[ABC^*][M]$$

$$\frac{d[ABC]}{dt} = k_3 \frac{k_1[A][BC]}{k_2 + k_3[M]} [M] = k'''_{complex} [A][BC]$$

$$k'''_{complex} = \frac{k_1 k_3 [M]}{k_2 + k_3 [M]}$$

At low pressure $k_2 > k_3[M]$

$$k'''_{complex,lowP} = \frac{K_1 K_3 [M]}{k_2}$$

$$k'''_{complex,lowP} = \frac{A_{3[} \exp(-\frac{T_3}{T})A_1 \exp(-\frac{T_1}{T})[M]}{A_2 \exp(-\frac{T_2}{T})} = \frac{A_1 A_3}{A_2} [M] \exp(-(\frac{T_3 + T_1 - T_2}{T}))$$

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$$k'''_{complex,lowP} = A \exp(-\frac{T_a}{T})[M]$$

$$T_i = \frac{E_a}{R}$$

At high pressure $k_3[M] > k_2$

$$k'''_{complex,highP} = k_1 = A_1 \exp(-\frac{T_1}{T})$$

Complex Reaction:

 $A + B \Leftrightarrow ABC \rightarrow AB + C$

- ABC is complex formed at top of energy hill.

- It referred to as transition state.

- Ter-molecular is just one good example of the complex elementary reaction, when the transition state of elementary reaction has a living long enough.



Distance along Reaction Coordinate

Chain reaction:

Initiation:
$$A_2 + h \nu \rightarrow 2A$$
 k_1 Propagation: $A + BC \rightarrow AB + C$ k_2 $AB + D \rightarrow DB + A$ k_3 Termination: $A + AB \rightarrow A_2B$ k_4

Chain Length: CL

$$CL_{A} = \frac{k_{2}[BC][A]}{k_{4}[A][AB]} = \frac{k_{2}[BC]}{k_{4}[AB]}$$

A and AB are chain carriers

$$\frac{1}{CL} = \frac{1}{CL_A} \frac{1}{CL_{BC}}$$

Chain reaction: Example

Initiation: $CH_4 + O_2 \rightarrow CH_3 + HO_2$

Propagation: 1st reaction

 $CH_3 + O_2 \rightarrow CH_2O + OH$

$$OH + CH_4 \rightarrow H_2O + CH_3$$

2nd reaction:

 $\begin{array}{l} OH+CH_2O\rightarrow H_2O+CHO\\ CHO+O_2\rightarrow HO_2+CO\\ HO_2+CH_2O\rightarrow H_2O_2+CHO \end{array}$

Intersection:

$$H_2O_2 + M \rightarrow 2OH + M$$

Propagation: 3rd reaction

$$OH + CO \rightarrow CO_2 + H$$

$$H + O_2 + M \rightarrow HO_2 + M$$
$$HO_2 + CO \rightarrow CO_2 + OH$$

Termination:

 $OH + HO_2 \rightarrow H_2O + O_2$ $HO_2 + HO_2 \rightarrow H_2O_2 + O_2$

Heterogeneous reactions

Homogeneous reactions: reactants are in the same phase, reactions between two gases, two liquids or two solids

Heterogeneous reactions: reactants are in two or more phases and/or reactions on the surface of a catalyst of a different phase

(manufacturing engineers have used surface reactions for synthesis of micro and nanoscale features in biomedical devices)



CO₂ formation

- The heterogeneous interaction of HCl with chlorine containing reservoir molecules (e.g. $CIONO_2$ and HOCl on low-temperature ices leads to the release of $Cl_2 \rightarrow$ effect on atmospheric chemistry in the polar regions ("ozone hole")
- heterogeneous reactions of NO_x leading to N₂O in the presence of SO₂ and airborne particles \rightarrow heavily polluted environments (power stations)



5 examples for heterogenous reactions:

• interaction of atmospherically relevant molecules (e.g. H₂O and NO₂ with combustion aerosol (soot) particles

 \rightarrow important in the planetary boundary layer and upper troposphere in relation to aircraft-induced Cirrus cloud formation

 NO_2 in a chemical oxidation-reduction reaction on soot result in HONO or in NO (depending on type of soot)

 \rightarrow HONO important OH source early in the morning which controls the development of photochemical smog

heterogeneous interaction of HX (X=CI, Br) on ice reveals that an interface region of approximately 100-200 nm thickness (depending on type of ice, the HX/ice interface controls a large extent the kinetics of effective bimolecular reactions with reservoir molecules e.g. CIONO₂ and HOCI)



Heterogeneous condensed surface phase reaction

Multiphase reactions, for example, SO_2 and H_2O_2 diffuse to the cloud droplet.

$SO_{2}(g) \xrightarrow{H_{2}O} SO_{2}(l)$ $SO_{2}(l) \xrightarrow{H_{2}O} HSO_{3}^{-} + H^{+}$ $HSO_{3}^{-} + H_{2}O_{2} \rightarrow SO_{4}^{2-} + H_{2}O + H^{+}$

In stratosphere,

$ClONO_2 + HCl(l) \rightarrow Cl_2 + HNO_3(l)$

It occurs in flog, aerosol, surface of ocean, plants or PSC.

Liquid phase reaction : diffusion controlled kinetic behaviour.

The diffusion-controlled rate constant (k_D) is $k_D = 4\pi r_{AB}^* D_{AB}$, where as D_{AB} is the diffusion coefficient for the reaction.

The example is Na⁺ in H₂O.
$$Na^+ + H_2O \rightarrow NaOH + H^+$$

$$D_{AB} = 1.3 \times 10^{-9} m^2 s^{-1}$$

 $r_{AB} = 0.5 nm$

$$k_{D} = 4\pi r_{AB} \cdot D_{AB} \cong 8 \times 10^{-12} \, cm^{3} \, molecule^{-1} s^{-1}$$
$$\cong 6.02 \times 10^{23} \cdot 8 \times 10^{-12} \cdot 10^{-3} \cong 5 \times 10^{9} \, dm^{3} \, mol^{-1} \cdot s^{-1}$$

In general, the rate constant of the gas $phase(k_G)$

 $k_G \equiv 2*10^{11} \text{ dm}^3 \text{ mol}^{-1} \text{ s}^{-1}$

Therefore, $k_D \equiv 400 k_G$ faster than gas phase reaction.



 $A + B \longrightarrow C$

Catalytic active surface

Liquid phase reaction



O¹D denotes electronically excited oxygen atom. A profit of O¹D can be found in NASA (1992).

Multiphase reactions

- 1. Gas diffusion
- 2. Absorption
- 3. Diffusion into the surface (not so fast)
- 4. Chemical reaction in liquid
- Liquid Diffusion of products 5. Gas Phase to the surface Gas phase diffusion Desorption of products 6. 6 Uptake Diffusion Bounce into bulk to the surface Evaporation off Reaction 5 7 Products Reaction at interface

Products

Uptake coefficient, γ , is the ratio of molecules lost to a surface to the number of gassurface collisions.

The rate of collision of a molecule X with an area A of the surface is ω then the rate of loss of X per unit volume (V) is equal to

$$\gamma \frac{\omega}{V}$$
$$\frac{d[X]}{dt} = \frac{\gamma \omega}{V} = \frac{\gamma c A[X]}{4V}$$

 $\overline{}$

therefore,

$$k_s = \frac{\gamma c A}{4V}$$

,

The partitioning of molecules between gaseous and liquid phases, the solubility of gas at low solute concentration obeys the Henry's law.

$$[X_s] = H_x P_x$$

where $[X_s]$ is the concentration of X in the solution, p_x is the pressure in the gas phase and H_x is the Henry's law coefficient.

The uptake coefficients at time t is

$$\frac{1}{\gamma_t} = \frac{1}{\gamma_0} + \frac{\pi^{\frac{1}{2}c}}{4H_x RTD^{\frac{1}{2}}} t^{\frac{1}{2}}$$

Consider reaction on the surface of atmospheric particles. The Langmuir adsorption isotherm is the simplest of equation to express the partitioning of gas between surface and gas phase. The surface coverage (θ_x) is ,

 $\theta_x = \frac{bp_x}{1+bp}$

where b is a constant equal to the ratio of rate coefficient for adsorption onto the surface and its desorption. The kinetics of reaction for a single reactant are developed by including a loss of the adsorbed molecule, the rate coefficient for this first-order loss process is k_1 . The chemical rate change is

$$Rate = k_L \theta_x = k_L b_x p_x$$







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Chemistry of the upper atmosphere

Energy source: Solar radiation, in short wavelength (UV-Vis) associate with the absorption of molecule by electronic transition.

Photolysis: Chemical changing and transition energy heat clumped into the atoms.

At the surface: absorption of visible energy leads to biological mechanism.

The first protection layer absorption are O, N, H- α band, N₂, it leads to the chemistry of ions and molecule in the heterosphere (ionosphere+ thermosphere, > 80 km).



The atmosphere is divided into 5 layers.

Troposphere is the first layer above the surface and contains half of the Earth's atmosphere. Weather occurs in this layer.

Stratosphere: Many jet aircrafts fly in this layer because it is very stable. Also, the ozone layer absorbs harmful rays from the Sun

Mesosphere: Meteors or rock fragments burn up in there; the atmosphere reaches its coldest temperature of around -90°C in the mesosphere.

Thermosphere layer with auroras, space shuttle orbits, the air is thin, small change in energy - large change in temperature (solar activity heat up to 1,500° C or higher)

Ionosphere an extension of the thermosphere, less than 0.1% of the total mass of the Earth's atmosphere (ionisation by solar radiation)

Exosphere: The region where atoms and molecules escape into space is referred to as the exosphere. The atmosphere merges into space in the extremely thin exosphere. This is the upper limit of our atmosphere.





Atmospheric phenomena characterizing in the Arctic middle atmosphere





The second protective layer is h_V (100 < λ < 200 nm). NO₂ source is transported from the troposphere, where is processed in soil (NO, N₂O)

$$N_2O + h\nu \rightarrow NO + N$$

$$N_2 O + h \nu \to N_2 + O$$

The third protective layer is the massive absorption from:

$$O_2(^3\sum_g^{-}) + h\nu \to O + O$$

In 1929, Chapman proposed the Chapman cycle:



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In 1929, Chapman proposed the Chapman cycle:

Production: $O_2 + h\nu \rightarrow O + O$ (1)

$$O + O_2 + M \to O_3 + M \tag{2}$$

Loss:

$$O + O_3 \to O_2 + O_2 \tag{3}$$

$$O_3 + h \nu \rightarrow O + O_2$$



Chapman-Cycle: explain qualitative the observed stratospheric O₃

(4)

- but: concentrations too high (factor 2), maximum too high
 - uncertainties in measurements can not explain the deviations

 \rightarrow loss process in addition ...

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In 1950-1960, using the Flash photolysis proved that k3 is slow, therefore it should be something going on in this region.

In 1940, Meinel observed the OH emission.

In 1952, Bates and Nicolet suggested odd oxygen cycles.

Cycle:1
$$H + O_3 \rightarrow OH + O_2$$
Cycle: 2
(35-60 km) $OH + O_3 \rightarrow HO_2 + O_2$ $\dot{OH} + O \rightarrow H + O_2$ $HO_2 + O \rightarrow OH + O_2$ Net: $O + O_3 \rightarrow O_2 + O_2$ Net: OH -source: $H_2O + h\nu \rightarrow OH + H$ H_2O -source in the stratosphere:
(Oxidation of CH4, not direct injection of
H_2O from troposphere)

The NO_x cycle: in 1979, aircraft emission from NO_x was found.

 $NO + O_3 \rightarrow NO_2 + O_2$ $NO_2 + O \rightarrow NO + O_2$



Net:
$$O + O_3 \rightarrow O_2 + O_2$$

<u>NO_x-source:</u> from active O(¹D) $O_3 + h\nu \rightarrow O(^1D) + O_2(^1\Delta)$ $O(^1D) + N_2O \rightarrow N_2 + O_2$ $O(^1D) + N_2O \rightarrow NO + NO$

from NO_x ion molecule down welling from the thermosphere and mesosphere.

- HO_x , NO_x and Chapman-cycle together agree with measured stratospheric O_3
- — Nobel prize 1995 (Paul Crutzen)

<u>NO₂ -source:</u> from troposphere => NO_3^- >> N_2O , N_2 , NO => NH_4^+

Loreloch measured the troposphere and found CFC's [CFC-11 (CF3CI) and CFC-12(CF2CI2)]. CFC is hanged around the atmosphere.

Rowland and Molina, in 1974, said CFC in the troposphere transports to the stratosphere.

In addition: CI and Br \rightarrow changing in the concentration of stratospheric O₃

<u>*ClO_x-cycle:*</u> (source: CFC-12: CF2Cl2 + $hv \rightarrow CF2Cl + Cl$)

 $\begin{array}{ll} \mathsf{CI} + \mathsf{O}_3 & \rightarrow \mathsf{CIO} + \mathsf{O}_2 \\ \\ \hline \mathsf{CIO} + \mathsf{O} & \rightarrow \mathsf{CI} + \mathsf{O}_2 \\ \hline \\ \hline \mathsf{net} \cdot \mathsf{O}_3 + \mathsf{O} & \rightarrow \mathsf{2O}_2 \end{array}$

Termination (formation of reservoirs) : $CI + CH_4 \rightarrow HCI + CH_3$ $CIO+ NO_2 + M \rightarrow CIONO_2 + M$

Re - release of CI : $HCI + OH \rightarrow CI + H_2O$ $CIONO_2 + h\nu \rightarrow CI + NO_3$



FIGURE 2-14

Ozone and chlorine monoxide concentrations versus latitude near the South Pole on September 16, 1987. (Source: Reprinted with permission from P. S. Zurer, *Chemical and Engineering News* (May 30, 1988): 16. Copyright 1988 by American Chemical Society.

1985: scientists analysed an O₃ decrease in polar regions



Figure 10-10 Vertical profiles of ozone over Antarctica measured by chemical sondes. In August the ozone hole has not developed yet, while in October it is fully developed. From Harris, N.R.P., et al., Ozone measurements, in WMO, op. cit..

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In upper and middle of the stratosphere, HO_x , NO_x , CIO_x , FO_x and BrO_x are involved.

Removal O₃ $X + O_3 \rightarrow XO + O_2$ k₁ Х $XO + O \rightarrow X + O_{\gamma}$ Life time of O_3 (k_{X+O3}[X]) k_2 Н $k_{H+O3}[H]$ OH $k_{OH+O3}[OH]$ $O + O_3 \rightarrow O_2 + O_2$ Net. NO k_{NO+O3}[NO] CI $k_{CI+O3}[CI]$ $\frac{d[X]}{dt} = -k_2[X][O_3] + k_2[XO][O] = 0$ k_{Br+O3}[Br] Br F $k_{F+O3}[F]$

Loss process

(x=H, OH, NO, Br and CI)

Source of OH and H (dominate in Mesosphere and upper stratosphere) $CH_4 + h\nu \to \to \to H_2O$ $H_2O + h\nu \to H + OH$

Source of NO_x (dominate in Stratosphere)

$$N_2O + O(^1D) \rightarrow NO + NO$$

 $(N_2O + hv \rightarrow N + NO)$

Source of Cl

$$O(^{1}D) + Cl - CF_{3} \rightarrow ClO + CF_{3}$$

CH₄ in Stratosphere oxidation

 $O_3 + h\upsilon \rightarrow O(^1D) + O_2$ $\rightarrow O(^{3}P) + O_{2}$ $O(^{1}D) + H_{2}O \rightarrow 2OH$ $O(^{1}D) + CH_{4} \rightarrow OH + CH_{3}$ $OH + CH_{\Lambda} \rightarrow H_2O + CH_3$ $O_2 + CH_3 + M \rightarrow CH_3O_2 + M$ $CH_{3}O_{2} + NO \rightarrow CH_{3}O + NO_{2}$ $CH_3O + O_2 \rightarrow HCHO + HO_2$ $HCHO + hv \rightarrow H + HCO$ or $OH + HCHO \rightarrow H_2O + HCO$ $O(^{1}D) + H_{2}O \rightarrow 2OH$ $HCO + O_2 \rightarrow HO_2 + CO$ $HO_2 + OH \rightarrow H_2O + O_2$ $CO + OH + O_2 \rightarrow CO_2 + HO_2$ $CH_4 + 2O_2 \rightarrow CO_2 + 2H_2O$ Net
How do they remove the ozone destroying? HOx

$$H_2 O \rightarrow H_2 + O$$

escape to the space

Termination

$$\begin{split} H, OH, HO_2 \Rightarrow OH + HO_2 \rightarrow H_2O + O_2 \\ OH + NO_2 + M \rightarrow HNO_3 + M \end{split}$$

 HNO_3 and H_2O are temporary reservoir of HO_x .

$$HNO_3 + h\nu \to H_2O + NO_2$$
$$H_2O + h\nu \to H + OH$$

Temporary reservoir of NO_x

 $NO_{2} + OH + M \rightarrow HNO_{3} + M$ at night: $NO_{2} + O_{3} \rightarrow NO_{3} + O_{2}$ $NO_{3} + hv \rightarrow NO_{2} + O$ $\rightarrow NO + O_{2}$ $NO_{3} + NO_{2} + M \rightarrow N_{2}O_{5} + M$ What remove the temporary reservoirs? $CIONO_{2} + hv \rightarrow CI + NO_{3}$

Temporary reservoir of CIO_x

 $CIO + NO_2 + M \rightarrow CIONO_2 + M$ $CI + CH_4 \rightarrow HCI + CH_3$ $HO_2 + CIO \rightarrow HOCI + O_2$

 $CIONO_{2} + h\upsilon \rightarrow CI + NO_{3}$ $\rightarrow CIO + NO_{2}$ $HCI + h\upsilon \rightarrow CI + H$

<u>CIOx</u>

Note: Highlight => propagation and termination



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$\underline{\text{BrO}_{x}} \text{ from } CFCl_{\underline{2}}Br \text{ (Halons)}$

In atmosphere, we have less Br than Cl. Br is heavy molecule and high vibrational energy.







Ozone is high produced in tropic region but amount of ozone is high in Poles.



- 1975: (IGY: International Geophysics Year) Dobson network is set for ozone measurement.
- 1985: Farmad proposed paper and we know the ozone hole. The explanation in the polar night

$$NO + O_3 \rightarrow NO_2 + O_2$$
$$NO_2 + O_3 \rightarrow NO_3 + O_2$$
$$NO_2 + NO_3 + M \Leftrightarrow N_2O_5 + M$$

In polar night, polar stratospheric clouds (PSCs) type I are formed.

$$N_2O_5(s) + H_2O(l) \rightarrow 2HNO_3$$

 $nHNO_3 + 3nH_2O \rightarrow n(HNO_3 \bullet 3H_2O)(NAT)$

Ozone destruction processes:

- 1. Winter polar vortex
- 2. $N_2O_5(s) \rightarrow NAT$
- 3. PSC formation (no photolysis)
- 4. Production of PSC product

 $CIONO_{2}(s) + HCI(s) \rightarrow CI_{2} + HNO_{3}$ $N_{2}O_{5} + HCI \rightarrow CIONO + HNO_{3}$ or $HOCI + HCI \rightarrow H_{2}O + CI_{2}$

5. During Spring

$CI_{2} + h\upsilon \rightarrow CI + CI$ $CIONO + h\upsilon \rightarrow CIO + NO$

6. Ozone destruction

$$2\{CI + O_3 \rightarrow CIO + O_2\}$$

$$CIO + CIO + M \rightarrow CI_2O_2 + M$$

$$CI_2O_2 + h\upsilon \rightarrow CI + CIOO$$

$$CIOO + M \rightarrow CI + O_2 + M$$

$$2O_3 \rightarrow 3O_2 \qquad Net$$

Stratospheric O₃

Polar stratospheric clouds (PSC) heterogenous reactions

$$\begin{split} & \text{CIONO}_{2} + \text{HCI} \xrightarrow{\text{PSC}} \text{CI}_{2} + \text{HNO}_{3} \\ & \text{CIONO}_{2} + \text{H}_{2}\text{O} \xrightarrow{\text{PSC}} \text{HOCI} + \text{HNO}_{3} \\ & \text{HOCI} + \text{HCI} \qquad \xrightarrow{\text{PSC}} \text{CI}_{2} + \text{H}_{2}\text{O} \\ & \text{N}_{2}\text{O}_{5} + \text{HCI} \qquad \xrightarrow{\text{PSC}} \text{CINO}_{2} + \text{HNO}_{3} \\ & \text{N}_{2}\text{O}_{5} + \text{H}_{2}\text{O} \qquad \xrightarrow{\text{PSC}} \text{2HNO}_{3} \end{split}$$

Result: CIO_x from reservoir and $NO_x \rightarrow HNO_3$ (**Denoxification**).

Stratosphere (dry) \rightarrow low temperatures (185K - 190K) are needed for condensation





6. März 2000 255 255 255

Ozone hole

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- UV light can have beneficial effects too
- It stimulates the transformation dietary steroids into Vit D (calciferol)
- Insects use it to guide them to nectar sources in flowers
- UV light is high energy light
- When it strikes molecules it can cause them to beak into ions or free radicals
- The free radicals in turn damage large molecules such as proteins and DNA - causes mutations - lead to cancer
 UV-B has a direct effect on DNA, DNA absorbs UV light of 260nm, the action of UV forms thymine dimers, this can cause a gene mutation when the DNA replicates





Arctic & Antarctic Trends



- Substantial losses have occurred in the Arctic
- Arctic levels are naturally higher because of the structure of the NH
 - 35-45 km, 30-50°N ozone
 - Note that the later year (post 1997) are above the downward slope
 - Ozone is no longer decreasing (1st step in the recovery process)

Newchurch et al. (2003)

Antarctic ozone hole theory

Solomon et al. (1986), Wofsy and McElroy (1986), and Crutzen and Arnold (1986) suggest reactions on cloud particle surfaces as mechanism for activating Chlorine



Cl₂ is easily photolyzed by UV & blue/green light HNO₃ is sequestered on PSC

Polar Ozone Destruction



Only visible light (blue/green) needed for photolyzing ClOOCl No oxygen atoms required Net: $2 \cdot O_3 +$ solar photon $\rightarrow 3 \cdot O_2$

Ozone Hole timeline



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How does chlorine get from our refrigerators to the Antarctic stratosphere?



Source Chemicals



- Cl is much more abundant than Br
- Br is about 50 times more effective at O₃ destruction

Atmospheric Chlorine Trends from NOAA/ERL – Climate Monitoring Division



Updated Figure made by Dr. James Elkins from Trends of the Commonly Used Halons Below Published by Butler et al. [1998], All CFC-113 from Steve Montzka (flasks by GC/MS), and recent updates of all other gases from Geoff Dutton (in situ GC).

Global Stratospheric Chlorine trends



Environmental Chemistry, A. Ladstätter-Weißenmayer WIC (2003) - FI

How do CFCs arrive the stratosphere?

- · CFCs have more weight than "air"
- · CFCs are stabil in the troposphere and therefore well mixed

from observations we know that CFCs arrive in the stratosphere but the concentration decrease with increasing height (photolysis)
 Atmospheric Measurements of CFC-11 and CF4



Source: CFC-12: $CF_2CI_2 + hv \rightarrow CF_2CI + CI$

 \rightarrow Montreal protocol and international appointments should prohibit or minimize the use of CFCs until 1996

Effect of the International Agreements on Ozone-Depleting Stratospheric Chlorine/Bromine









Correlation of ozone concentration - global warming?

Ozone absorbs UV-radiation, increase of temperature in stratosphere \rightarrow less ozone lower temperature in stratosphere

Greenhouse gas absorbs IR radiation from the surface, increase of temperature in lower atmosphere but cooling of upper atmosphere \rightarrow global warming leads to cooler stratosphere

Reduction of stratospheric temperature, increase of PSC a in spring reduction of ozone, increase of heterogenous processes in mid-latudues possible

What about the future of ozone reduction?

Global reduction -5% since 1965

10% reduction in mid-latitude winter/spring

60% reduction in antarctic spring

5-30% reduction in arctic spring

up to 20% reduction after volcanic eruption of Pinatubo (Philippines)

Maximum of stratospheric CI now reached ?

back to "normal situation" in 2100?

1st detection is the approximate year that we'll be able to say that the ozone hole is improving.

Full recovery is the year when the ozone hole size is zero.



•

Ozone and climate change



Ozone change is not a primary cause of climate change.

- Ozone depleting substances contribute to climate change.
- Ozone changes causes a climate response that is generally larger than the Halocarbon response.
- Climate change may seriously impact ozone levels.

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Stratosphere	Troposphere
Low pressure	High pressure
Low temperature	High temperature
Some sinks	Many sinks
Some sources	Many sources
No biological emissions	Biological emissions
Less anthropopgenic input	Many anthropopgenic input
Less vertical mxing	More vertical mixing
Low humidity	High humidity

Emissions in Troposphere



- Many different trace gases in troposphere, Complex reaction chains
- Life times from seconds to years, Transport from source to sink

Life time of trace gases

- source \rightarrow sink
- dry deposition: gases are absorbed
- irreversible at the surface, Ocean and plant-surfaces
- wet deposition: compounds are implemented in clouds, aerosols, rain droplets
- emit and remove from atmosphere
- Transport: vertical und horizontal
- Source \rightarrow Sink = Life time

Mixing time: T_M

Vertical mixing: ~ 1 week (surface \rightarrow tropopause)

Horizontal mixing:~ 1 year (around the globe)

e.g. CH_4 (by physical processes and oxidation \rightarrow slowly oxidydised τ = some years), HCHO (intermediate during oxidation, fast photolysed τ = some hours)

- $H_2 \tau$ = some years CO τ = some months
- natural gases: biogenic sources, reduced form (CH₄, terpenes, H₂S, ...) or only oxidised (CO, N₂O) = mainly solute in water

"anthropogenic gases": higher oxidised (CO₂, NO₂, SO₂) easy soluble

• only some are directly deposited

Dry deposition = SO_2 , O_3 , CO_2 , SO_3

- Micro biological sinks = CO, H_2
- fast soluble = HCl, HF, HNO_3
- Wet deposition, after arriving in aerosols = NH_3 , SO_2 , NO_x
- Inert (physical and chemical) = NO_x , Carbonyl, Sulfids (COS), CFCs
- all others = photochemical reactions with OH, or direct photolysis (O_3 , NO_2 , HCHO)
- near the surface: many turbulences, mixing (0.5-2 km during day and less during night)
 - time scale:

days:	some compounds, atmospheric circulation, transport
Free troposphere:	transport over long distances near surface
Wind speed:	10-30 ms ⁻¹ in E-W direction, less in N-S and vertical direction (mid latitude)

Air masses are moving in circles along latitudes in 10s of days

Movings in N-S directions slower, transport between hemispheres not allowed, only during defined seasons (N \rightarrow S in upper troposphere and S \rightarrow N in lower troposphere (1 year)

E-W averaging (tropospheric mixing $\tau = 1$ month (2-dimensional model (height, latitude)

Many compounds:

- 1940: ~24 different chemical compounds were known
- 1950: known scientists begun the resarch on chemical and biological compounds of atmoshere, less than 100 compounds were known
- today: 3000 compounds are known
- some in atmospheric gases, aerosol particles, (small and big), in rain snow, fog, in reservoirs
- natural and anthropogenic sources, pre-industrial atmosphere

more or less reactive compounds ($CH_4 = 1.7-1.8 \text{ ppmv}$)

4 times more reactive are organic than inorganic compounds

Many compounds:

.

- organic: HC (solutions, industrial chemicals, combustion processes)
- reaction to oxidised products (moving faster than precursors)
- every element of periodic system is in atmosphere
- a small number of groups of compounds:
- sulfur compounds
- nitrogen compounds
- HC compounds
- Halogen compounds
- with and without HC

Emission of compounds which are removed \rightarrow cycles of elements (*bio geo chemical cycles: e.g. B. H*₂O)

Transport: within the atmosphere, ocean, biosphere, continent

• "Elements of life": C, O, N, S, P with reservoir substances, within the atmosphere, ocean, organism

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- many source gases partiell oxidized (CO, SO_2) or reduced (H_2S , NH_3)
- in the troposphere they are not directly oxidized by O₂
- first reaction is with hydroxy radical **OH** (atmospheric cleaning substance)
- <u>example:</u>
- $CH_4 + 2O_2 \rightarrow CO_2 + 2H_2O$
- Oxidation follows many reaction steps and produce redicals as function of photolysis (OH-concentration, other radicals and NO_x).

- OH is the most important radical in troposphere
- many tropospheric reactions start with OH (life time_{OH} ~ 1 s, low concentration ~10⁶ molec/cm³)
- concentration variable dependent of: photolysis, O₃- and H₂O-concentration and other radicals

Sources of OH:

 $O(^{1}D) + H_{2}O \rightarrow 2OH$

 $\mathrm{O_3} + \mathrm{hv} \rightarrow \mathrm{O_2} + \mathrm{O(^1D)}$

Photolysis of HCHO:

$$CH_2O + hv \rightarrow H + HCO$$
$$\rightarrow H_2 + CO$$
$$HCO + O_2 \rightarrow HO_2 + CO$$
$$H + O_2 + M \rightarrow HO_2 + M$$

under clean air situation no NO:

 $\mathrm{HO}_2 + \mathrm{O}_3 \longrightarrow \mathrm{OH} + \mathrm{2O}_2$

 $HO_2 + NO \rightarrow OH + NO_2$

- 2. Reaction of O_3 with alkenes
- 3. Reaction of NO_3 with aldehyds or alkens during night

- <u>Sinks for OH:</u>
- 1. Direct key reactions:

 $HO_2 + HO_2 \rightarrow H_2O_2 + O_2$ $OH + HO_2 \rightarrow H_2O + O_2$

2. Formation of stabil peroxiyds

 $\rm CH_3O_2 + HO_2 \rightarrow \rm CH_3OOH + O_2$

3. Reaction with $NO_2 \& HNO_3$ in polluted areas

 $OH + NO_2 + M \rightarrow HNO_3 + M$ $OH + HNO_3 \rightarrow H_2O + NO_3$

A: Reaction with OH

.

$SO_2 + OH$	\rightarrow OSOOH
CO + OH	\rightarrow HOCO
$CH_4 + OH$	$\rightarrow CH_3 + H_2O$
NH ₃ + OH	$\rightarrow \mathrm{NH_2} + \mathrm{H_2O}$
$H_2S + OH$	\rightarrow SH + H ₂ O
CH ₃ CI + OH	$\rightarrow CH_2CI + H_2O$

B: Photolysis

HCHO + hv \rightarrow H + HCO (λ < 338 nm)

C: Reaction of radicals with O₂

 $CH_3 + O_2 \rightarrow CH_3OO$

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D: Reaction of Peroxy-Radicals with NO

 $HOO + NO \rightarrow OH + NO_2$

 $\rm CH_3OO + \rm NO \rightarrow \rm CH_3O + \rm NO_2$

E: Abstraction of H by O₂:

 $CH_3O + O_2 \rightarrow CH_2O + HO_2$

 $\rm HOCO + O_2 \rightarrow \rm OCO + \rm HO_2$

 $\rm HCO + O_2 \rightarrow \rm CO + \rm HO_2$

F: Addition von O₂:

 $\mathsf{RCO} + \mathsf{O_2} \to \mathsf{ROCOO}$



$$NMHC + NO + hv \rightarrow NO_2 + \Pr{odukte}$$
 R1

$$NO_2 + h\nu(\lambda \le 410nm) \rightarrow NO + O$$
 R2

$$O + O_2 + M \to O_3 + M$$
 R3

$$NO + O_3 \rightarrow NO_2 + O_2$$
 R4

equilibrium:

$$\frac{[NO_2]}{[NO]} = \frac{k_{R4}[O_3]}{j_{R2}}$$

 $O_{3} + hv (\leq 310nm) \rightarrow O(^{1}D) + O_{2}$ $O(^{1}D) + H_{2}O \rightarrow OH + OH$ $RH + OH \rightarrow R + H_{2}O$

 $R + O_2 + M \rightarrow RO_2 + M$

 $RO_2 + NO \rightarrow NO_2 + RO$

Methan-Oxidation $O_3 + h\nu \rightarrow O(^1D) + O_2$ wavelenght: < 310 nm</td> $O(^1D) + H_2O \rightarrow OH + OH$ $O(^1D) + M \rightarrow O + M$ $O(^1D) + M \rightarrow O + M$ Regeneration of O_3

Oxidation steps:

 $OH + CO \rightarrow H + CO_{2}$ $OH + CH_{4} \rightarrow CH_{3} + H_{2}O$ $H + O_{2} \rightarrow HO_{2}$ $CH_{3} + O_{2} \rightarrow CH_{3}O_{2}$

Methan-Oxidation

 $CH_3O + O_2 \rightarrow HCHO + HO_2$ $HCHO + h\nu(\lambda \leq 330nm) \rightarrow H + HCO$ $HCO + O_2 \rightarrow CO + HO_2$ $H + O_2 + M \rightarrow HO_2 + M$ $3[H + O_2 + M \rightarrow HO_2 + M]$ $4[NO_2 + h\nu(\lambda \le 410nm) \rightarrow NO + O]$ $3[O+O_2+M \rightarrow O_3+M]$ $CH_4 + 8O_2 + 5h\nu \rightarrow CO + 4O_3 + 2OH + H_2O$ Net:

CO-Oxidation



Ozon-Production in Troposphere

 $\begin{aligned} NO_2 + h \nu &\to O + NO & \text{high NO}_x - \text{production of O}_3: \\ O + O_2 + M &\to O_3 + M \\ OH + CO &\to H + CO_2 \\ H + O_2 + M &\to HO_2 + M \\ HO_2 + NO &\to OH + NO_2 \end{aligned}$

Netto: $CO + 2O_2 + h\nu \rightarrow CO_2 + O_3$

Ozon-Destruktion in der Troposphäre

 $OH + CO \rightarrow H + CO_2$ low NO_x – destruction of O₃: $H + O_2 + M \rightarrow HO_2 + M$ $HO_2 + O_3 \rightarrow OH + 2O_2$

Netto: $CO + O_3 \rightarrow CO_2 + O_2$

Methan-Oxidation

Production of peroxy radicals:

 $H + O_2 + M \rightarrow HO_2 + M$

 $CH_3 + O_2 + M \rightarrow CH_3O_2 + M$

low NO_x – production of acids:

 $HO_{2} + HO_{2} \rightarrow H_{2}O_{2} + O_{2}$ $CH_{3}O_{2} + HO_{2} \rightarrow CH_{3}OOH + O_{2}$ $high NO_{x} - production of NO_{2}:$ $HO_{2} + NO \rightarrow OH + NO_{2}$ Regen

 $CH_3O_2 + NO \rightarrow CH_3O + NO_2$

Regeneration of OH:

Or production of Methoxy-Radical:

CO, CO₂, und CH₄ Messungen von SCIAMACHY

CO column [10¹⁸/cm²]

4.00

3.70

3.40

3.10

2.80 2.50 2.20

1.90

1.60

1.30

1.00



XCO2 [ppmv] Feb 385 380 375 370 May - Jul 365 360 355 350 Aug - Oct 345 340

Buchwitz et al., de Beek et al., ACPD, 2005, ACPD 2006

Carbon monoxide SCIAMACHY 2003



Carbon dioxide SCIAMACHY 2003 335

CH₄ Messungen von SCIAMACHY



Environmental Chemistry, A. Ladstätter-Weißenmayer



A. Richter et al., Increase in tropospheric nitrogen dioxide over China observed from space, Nature, 437 2005



Courtesy of Andreas Richter

HCHO Messungen von GOME



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Troposphärische O₃ Messungen von GOME





Environmental Chemistry, A. Ladstätter-Weißenmayer

Anorganic Chemistry

 $HSO_{3}^{-} + H_{2}O_{2} \rightarrow HSO_{4}^{-} + H_{2}O$ $HSO_{3}^{-} + O_{3} \rightarrow HSO_{4}^{-} + O_{2}$ $H_{2}SO_{4} \leftrightarrow H^{+} + HSO_{4}^{-}$ $HSO_{4}^{-} \leftrightarrow H^{+} + SO_{4}^{2-}$ $HNO_{3} \leftrightarrow H^{+} + NO_{3}^{-}$

Photochemistry – Organic Chemistry

 $H_2O_2 + hv(\lambda \le 380nm) \rightarrow 2OH$

 $RCHO_{(aq)} + OH \rightarrow RC(O)OH_{(aq)}$

 $RC(O)OH \leftrightarrow H^+ + RC(O)O^-$

- · compounds caused by antropogenic activities
- concentrations high above normal
 - \rightarrow human influence, animals, veetation and material
- mainly: in urban and industrialised regions
- 1. Typ: (SO₂ and Sulfate), particles from combustion
- cold regions: heating Emissions
 - 2. Gas and Benzene (Los Angeles, Tokyo, Athens, Mexico City, Sao Paulo, Brazil)









1952: London view less than some m for 1 week



Smog (photochemical Smog)

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Smog: NO_x , organic compounds, O_3 (nitrats, oxidised HC, photochemical aerosols)

O₃ normal: 20-60 ppb

O₃ urban: 500 ppb

Natural and anthropogenic processes (started mainly with OH)

High reactivity \rightarrow Oxidation and chemische reaction

Free Radicals - chain reactions: Oxidation of $\rm H_2, \ CH_4, \ other \ HC, \ CO \ to \ CO_2 \ and \ H_2O$

other gases like NO_{x} and sulfur compouds modify the "way of combustion process"

surface as source of compounds

 NO_{x} and CO produced in sand storm surface emission: natural and anthropogenic sources

High concentration in urban and industrial regions (NH)

Stratospheric O₃ limits sun radiation at surface \rightarrow wavelenght > 280 nm

Photochemical labill > 280 nm: O_3 , NO_2 , HCHO

All 3 gases (O₃, NO₂, HCHO) \rightarrow OH (oder HO₂ production) \rightarrow start of oxidation chain

- 10% of total O₃ in troposphere
- natural atmosphere: all reactions start with O₃
- Source of O₃ is important

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- O_3 formation in low altitude (dependent of hydrocarbons and NO_x)
- O₃ Transport (Stratospheric Tropospheric Exchange)

- OH influence during day (stratosphere and troposphere)
- NO_3 influence during night
- NO_3 less reactiv than OH
- · diurnal variation different (OH and NO_3)
- OH photolysed (during day)
- NO_3 photolysed during day (present during night)
- 90% of total atmospheric mass stays in troposphere (main part of gases of combustion processes)



IPCC 2007



IPCC 2007

since 1750: CH₄- and N₂O- increase ~148% respectively ~ 18 %



- GHG: H₂O, H₂, CO₂, O₃, N₂O, CH₄.
- important for climate change caused by physical and chemical features
 - GHG are gases in an atmosphere that absorb and emit radiation within the thermal infrared range. This process is the fundamental cause of the greenhouse effekt
 - GHG warm the atmosphere by efficiently absorbing thermal infrared radiation emitted by the Earth's surface, by the atmosphere itself, and by clouds. As a result of its warmth, the atmosphere also radiates thermal infrared in all directions, including downward to the Earth's surface. Thus, greenhouse gases trap heat within the surface-troposphere system.

- in addition anthropogenic synthetic gases (halogen compounds) are emitted \rightarrow climte change
- the measured value: "Global Warming Potential" (efficiency of a gas over 100 years is observed, fficiency of CH_4 is 21 time higher than CO_2)
- if the concentration of emitted gases is known \rightarrow Global Warming Potential can be calculated
- More GHG more climate change
Tropospheric Chemistry



Source: Mann et al. 1999.



